

## Governing Equations in Computational Fluid Dynamics: Derivations and A Recent Review

Review  
Article

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### ABSTRACT

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The objective of this paper is to provide quick, complete and up-to-date reference on governing equations applied in computational fluid dynamics (CFD) related research, along with the recent review on their future development. The development of non-Newtonian momentum equations, formation of conservation equations in advanced coordinate systems and inclusion of miscellaneous body forces into momentum equations are highlighted. This may ease complicated numerical burdens in solving fluid dynamics equations. Continuity, Navier-Stokes and energy equations are involved, while their coordinate systems span across Cartesian, cylindrical and spherical domain.

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## 1. Introduction

Continuity equations, Navier-Stokes equations and energy equations are the key governing equations that dictate the physics of fluid mechanics and thermal sciences, which are instrumental in the research of computational fluid dynamics (CFD). From simple creeping flow [1,2] and Couette flow [3,4], to the state-of-the-art turbulence modelling [5,6], moving boundaries simulation [7-9], nanofluids motion [10,11], multiphase flow [12,13], complex geometry aerodynamic design [14,15], wave modelling [16,17] and oceanic engineering [18,19], all these engineering researches fall under the purview of these governing equations.

The governing equations stem from the fundamental principles of Newton's Laws and Reynold's transport theorem [20,21], which can be expressed in a general form of integral equations. However, such a general form is not convenient for the precise analysis down to the scale of fluid element parcels [22]. Eulerian approach therefore takes the stage, and it is further developed into the differential form of equations which involve tensors and indicial notation for spatial description and flow fields. Introduction of continuum mechanics and constitutive laws [23,24] had laid down the cornerstone for the derivation of these conservation equations into the governing equations as what we see today.

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The general description on the formula in CFD could be found in many textbooks [25-27], yet there is a lack of detailed derivations in a single work with recent reviews. Hence the paper integrates the derivations of these equations with its recent reviews on the essential future development. Section 2 will address the conservative equations, while in Section 3 the development and future research on the conservative equations are highlighted. The development of non-Newtonian momentum equations, formation of conservation equations in advanced coordinate systems and inclusion of miscellaneous body forces into momentum equations are discussed.

## 2. Governing Equations in CFD

Consider an infinitesimal fluid elements as shown in **Fig. 1 - 3**, which represents the flow field domain based on Cartesian, cylindrical and spherical coordinate respectively. The term  $\kappa_{S1,S2}$  is a general representation of flow field vector, in which the subscript S1 and S2 is the spatial component for the flow field vector.

### 2.1 Continuity Equations

Continuity equation can be perceived as nonlinear diffusion equation with regular drift term, and it inspires ubiquitous applications in many fields such as crowd modelling [28], prediction of aerospace debris cloud evolution [29], biomedical imaging [30] and curve measurement analysis [31]. The equation can be treated as either initial boundary problem [32] or Cauchy problem [33] too.

The fundamental physics of Continuity Equations is the principle of conservation of mass, proposed by Lavoisier [34] in 1985. Conservation of mass can be defined as: the conservation law that the rate of change of mass within a control volume (CV) is equivalent to the net rate of mass flowing into the CV [35,36]. Consider the integral form of the mass conservation equation:

$$\frac{\partial}{\partial t} \int_{CV} \rho \cdot dV + \int_{CS} \rho \mathbf{v} \cdot \mathbf{n} \cdot dA = 0, \quad \forall \mathbf{v} \in \mathbb{R} \quad (1)$$

Eq. (1) can be transformed to differential form using Gauss' divergence theorem [20,21] to form:

$$\dot{\rho} + \nabla \cdot (\rho \mathbf{v}) = 0 \quad (2)$$

However, the paper will demonstrate the derivation approach based on the CV facet analysis.  $\nabla$  is the divergence term which can be defined based on its coordinate system, which can be further developed into Eq. (5), (8) or (11).

#### 2.1.1 Cartesian coordinate Continuity equation

Consider **Fig. 1**, the length of the infinitesimal fluid element in  $x$ ,  $y$ , and  $z$  direction can be assigned as  $\delta x$ ,  $\delta y$  and  $\delta z$  respectively. The term  $\kappa_{S1,S2}$  in **Fig. 1** can be defined as:

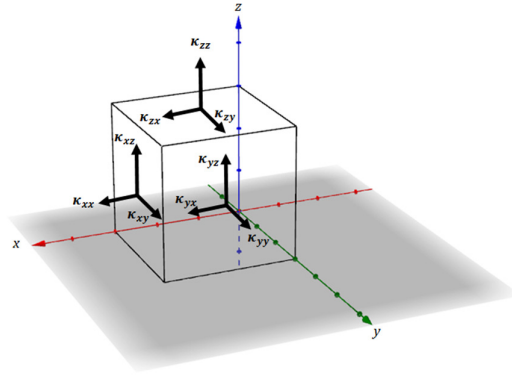
$$\kappa_{S1,S2} = \langle \kappa_{x,x} \quad \kappa_{y,y} \quad \kappa_{z,z} \rangle = \left\{ \left\langle \frac{\partial(\rho u_x)}{\partial x} \quad \frac{\partial(\rho u_y)}{\partial y} \quad \frac{\partial(\rho u_z)}{\partial z} \right\rangle \cdot V_{car} \mid V_{car} = \delta x \delta y \delta z \right\} \quad (3)$$

The other  $\kappa$  are zero due to the non-slip boundary condition and by substituting Eq. (3) into Eq. (1),

$$\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u_x)}{\partial x} + \frac{\partial(\rho u_y)}{\partial y} + \frac{\partial(\rho u_z)}{\partial z} = 0 \quad (4)$$

Taking  $\mathbf{v} = [u_x \quad u_y \quad u_z]$ , in an incompressible flow, Eq. (3) will be reduced to  $\nabla \cdot \mathbf{v} = 0$  where:

$$\nabla = \frac{\partial}{\partial x} + \frac{\partial}{\partial y} + \frac{\partial}{\partial z} \quad (5)$$



**Fig. 1.** Infinitesimal fluid field domain based on Cartesian coordinate

### 2.1.2 Cylindrical coordinate Continuity equation

Consider **Fig. 2**, the length of the infinitesimal fluid element in  $r$ ,  $\theta$ , and  $z$  direction can be assigned as  $\delta r$ ,  $\delta\theta$  and  $\delta z$  respectively. Upon dimensional expansion, these distances will evolve as  $r+\delta r$ ,  $\theta+\delta\theta$  and  $z+\delta z$  respectively. The term  $\kappa_{S1,S2}$  in **Fig. 2** can be further defined as:

$$\kappa_{S1,S2} = \langle \kappa_{r,r} \quad \kappa_{\theta,\theta} \quad \kappa_{z,z} \rangle = \left\{ \left( \frac{\partial(\rho u_r)}{\partial r} + \frac{\rho u_r}{r} \right) \frac{1}{r} \frac{\partial(\rho u_\theta)}{\partial \theta} + \frac{\partial(\rho u_z)}{\partial z} \right\} \cdot V_{\text{cyn}} \Big| V_{\text{cyn}} \approx r \delta\theta \delta r \delta z \quad (6)$$

The value of other non-normal  $\kappa$  is zero too due to the non-slip boundary condition. Note that  $\delta r^2 \approx 0$  during the derivation due to its infinite proximity to zero. The volume of the cylinder is:

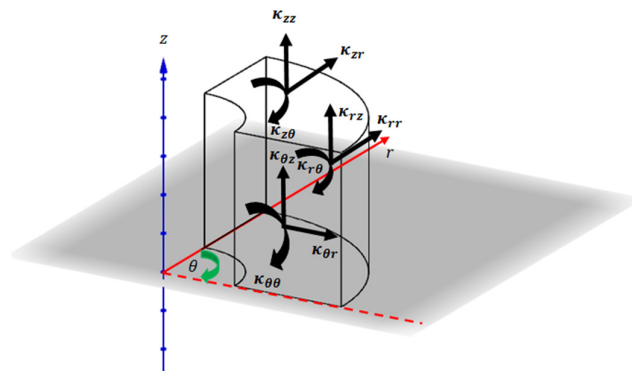
$$V_{\text{cy}} = (\pi(r+\delta r)^2 - \pi r^2) \times \frac{\delta\theta}{2\pi} \times \delta z \approx r \delta\theta \delta r \delta z \quad \blacksquare$$

Substitute Eq. (6) into Eq. (1) will yield:

$$\frac{\partial \rho}{\partial t} + \frac{\rho u_r}{r} + \frac{\partial(\rho u_r)}{\partial r} + \frac{1}{r} \frac{\partial(\rho u_\theta)}{\partial \theta} + \frac{\partial(\rho u_z)}{\partial z} = 0 \quad (7)$$

If the flow is incompressible, Eq. (7) can be simplified into  $\nabla \cdot \mathbf{v} = 0$  too with the divergence term as in Eq. (8), provided that the velocity vector is  $\mathbf{v} = [u_r \quad u_\theta \quad u_z]$ .

$$\nabla = \frac{1}{r} \frac{\partial}{\partial r} + \frac{1}{r} \frac{\partial}{\partial \theta} + \frac{\partial}{\partial z} \quad (8)$$



**Fig. 2.** Infinitesimal fluid field domain based on cylindrical coordinate

### 2.1.3 Spherical coordinate Continuity equation

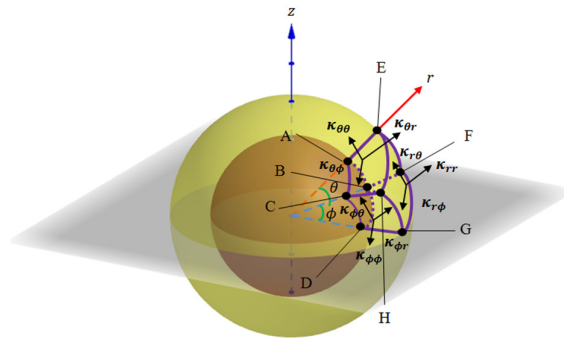
From **Fig. 3**, the length of control volume boundary and surface of the facets for the spherical fluid element can be shown in **Table 1**. All the  $\kappa$  will be zero as well, except  $\kappa_{r,r}$ ,  $\kappa_{\theta,\theta}$  and  $\kappa_{\phi,\phi}$  which can be mathematically expressed as in Eq. (9).

$$\begin{aligned} \kappa_{S1,S2} &= \langle \kappa_{r,r} \quad \kappa_{\theta,\theta} \quad \kappa_{z,z} \rangle \\ &= \left\langle \left( \rho u_r + \frac{\partial(\rho u_r)}{\partial r} \delta r \right) A_{EFGH} - (\rho u_r) A_{ABCD} \quad \left( \rho u_\theta + \frac{\partial(\rho u_\theta)}{\partial \theta} \delta \theta \right) \cdot A_{BDFG} - (\rho u_\theta) \cdot A_{ACEH} \quad \left( \rho u_\phi + \frac{\partial(\rho u_\phi)}{\partial \phi} \delta \phi \right) \cdot A_{CDGH} - (\rho u_\phi) \cdot A_{ABEF} \right\rangle \\ &= \left\{ \left( \frac{\partial(\rho u_r)}{\partial r} + \frac{2\rho u_r}{r} + \frac{1}{r} \frac{\partial(\rho u_\theta)}{\partial \theta} + \frac{\rho u_\theta}{r} \frac{\partial(\rho u_\phi)}{\partial \phi} \right) \cdot V_{\text{sph}} \Big| V_{\text{sph}} \approx r^2 \sin(\theta) \delta r \delta \theta \delta \phi \right\} \end{aligned} \quad (9)$$

The volume of spherical element can be approximated by taking the product of  $A_{ACEH}$  and  $L_{CD}^*$ , or using the Jacobian rules [37] for the derivation. Substitute Eq. (9) into Eq. (1) will form the compressible Continuity equation as in Eq. (10), in which  $\nabla \cdot \mathbf{v} = 0$  where  $\mathbf{v} = [u_r \quad u_\theta \quad u_\phi]$  will be applied in incompressible case where its divergence term is shown in Eq. (11).

$$\frac{\partial \rho}{\partial t} + \frac{2\rho u_r}{r} + \frac{\partial(\rho u_r)}{\partial r} + \frac{\rho u_\theta}{r} \cot(\theta) + \frac{1}{r} \frac{\partial(\rho u_\theta)}{\partial \theta} + \frac{1}{r \sin(\theta)} \frac{\partial(\rho u_\phi)}{\partial \phi} = 0 \quad (10)$$

$$\nabla = \frac{1}{r^2} \frac{\partial(r^2)}{\partial r} + \frac{1}{r \sin(\theta)} \frac{\partial(\sin \theta)}{\partial \theta} + \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} \quad (11)$$



**Fig. 3.** Infinitesimal fluid field domain based on spherical coordinate

**Table 1** Geometry analysis on spherical fluid element based on **Fig. 3**

Geometry	Equation
LAB	$r \delta \theta$
LBC	$r \sin(\theta + \delta \theta) \delta \phi$
LCD	$r \delta \theta$
LAD	$r \sin(\theta) \delta \phi$
LEF	$(r + \delta r) \delta \theta$
LFG	$(r + \delta r) \sin(\theta + \delta \theta) \delta \phi$
LGH	$(r + \delta r) \delta \theta$
LEH	$(r + \delta r) \sin(\theta) \delta \phi$
L <sub>AE</sub> , L <sub>BF</sub> , L <sub>CG</sub> , L <sub>DH</sub>	$\delta r$
A <sub>ABCD</sub>	$r^2 \sin(\theta) \delta \theta \delta \phi$
A <sub>AEFGH</sub>	$(r + \delta r)^2 \sin(\theta) \delta \theta \delta \phi$
A <sub>ABDFG</sub>	$r \sin(\theta + \delta \theta) \delta r \delta \phi$
A <sub>ACEH</sub>	$r \sin(\theta) \delta r \delta \phi$
A <sub>ABEF</sub> , A <sub>CDGH</sub>	$r \delta \theta \delta r$

\* $\delta r^2 = \delta r^3 \approx 0$ ,  $\sin(\delta \theta) \approx \delta \theta$ ,  $\cos(\delta \theta) \approx 1$

## 2.2 Navier-Stokes Equations

Momentum equations are originated from Newton's second law which states that force of a moving object is equivalent to its rate of change of momentum. Expanding the definition will give the momentum equations in a general integral form of [35,38]:

$$\sum F = \frac{\partial}{\partial t} \int_{CV} \rho \mathbf{v} \cdot dV + \int_{CS} \rho (\mathbf{v}\mathbf{v}) \cdot \vec{n} \cdot dA, \quad \forall \mathbf{v} \in \mathbb{R} \quad (12)$$

The first term of Eq. (12) represents the body forces which may include gravity, Coriolis effects, centrifugal force and electromagnetic force [36]; while the second term denotes the surface forces, which typically refers to pressure force and viscous force. If the flow is in steady state, then  $\sum F$  will be negated. Eq. (12) was expanded by French mathematician Augustin Louis de Cauchy [39] into differential term with the application of divergence theorem in such a way that:

$$\frac{\partial}{\partial t} (\rho \mathbf{v}) + \nabla \cdot (\rho \mathbf{v}\mathbf{v}) = \rho \mathbf{g} + \nabla \cdot \vec{\sigma}_{ij} \quad (13)$$

$$\frac{\partial}{\partial t} (\rho \mathbf{v}) + \nabla \cdot (\rho \mathbf{v}\mathbf{v}) = \rho \left( \frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \right) = \rho \frac{D\mathbf{v}}{Dt} \quad (14)$$

Eq. (14) is indeed the material derivative [21], and sometimes it is named as total, particle, Lagrangian, Eulerian or substantial derivatives [26]. It stands for the convection phenomenon and its cancellation implies the formation of creeping flow.  $\rho \mathbf{g}$  and  $\nabla \cdot \vec{\sigma}_{ij}$  represents body force and sum of applied surface forces, respectively [21]. However, the exact equation for divergence term  $\nabla$  will be varied from the coordinate systems. Eq. (13) needs to be further developed, and it will evolve as the famous Navier-Stokes equations [40]. This implies that only Newtonian fluid is considered, while the non-proportional relationship between velocity field and stress tensor which exists in non-Newtonian fluid requires additional modelling [41-44], and it will not be covered in this section.

### 2.2.1 Cartesian coordinate Navier-Stokes Equations

The material derivative of the Cartesian coordinate can be expanded as:

$$\frac{D\mathbf{v}}{Dt} = \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{v}}{\partial x} \frac{\partial x}{\partial t} + \frac{\partial \mathbf{v}}{\partial y} \frac{\partial y}{\partial t} + \frac{\partial \mathbf{v}}{\partial z} \frac{\partial z}{\partial t} = \frac{\partial \mathbf{v}}{\partial t} + u_x \frac{\partial \mathbf{v}}{\partial x} + u_y \frac{\partial \mathbf{v}}{\partial y} + u_z \frac{\partial \mathbf{v}}{\partial z} \quad (15)$$

The  $\kappa_{S1,S2}$  term in this section refers to viscous forces acting on the control surface of fluid element per unit volume:

$$\boldsymbol{\kappa} = \begin{pmatrix} \kappa_{x,x} & \kappa_{x,y} & \kappa_{x,z} \\ \kappa_{y,x} & \kappa_{y,y} & \kappa_{y,z} \\ \kappa_{z,x} & \kappa_{z,y} & \kappa_{z,z} \end{pmatrix} = \begin{pmatrix} \frac{\partial \tau_{xx}}{\partial x} & \frac{\partial \tau_{xy}}{\partial x} & \frac{\partial \tau_{xz}}{\partial x} \\ \frac{\partial \tau_{yx}}{\partial y} & \frac{\partial \tau_{yy}}{\partial y} & \frac{\partial \tau_{yz}}{\partial y} \\ \frac{\partial \tau_{zx}}{\partial z} & \frac{\partial \tau_{zy}}{\partial z} & \frac{\partial \tau_{zz}}{\partial z} \end{pmatrix} = \frac{\partial \boldsymbol{\tau}}{\partial \mathbf{X}} \quad (16)$$

$$\boldsymbol{\tau} = \begin{pmatrix} \tau_{xx} & \tau_{xy} & \tau_{xz} \\ \tau_{yx} & \tau_{yy} & \tau_{yz} \\ \tau_{zx} & \tau_{zy} & \tau_{zz} \end{pmatrix} = \begin{pmatrix} 2\mu \frac{\partial u_x}{\partial x} + \lambda \nabla \cdot \mathbf{v} & \mu \left( \frac{\partial u_x}{\partial y} + \frac{\partial u_y}{\partial x} \right) & \mu \left( \frac{\partial u_x}{\partial z} + \frac{\partial u_z}{\partial x} \right) \\ \mu \left( \frac{\partial u_y}{\partial x} + \frac{\partial u_x}{\partial y} \right) & 2\mu \frac{\partial u_y}{\partial y} + \lambda \nabla \cdot \mathbf{v} & \mu \left( \frac{\partial u_y}{\partial z} + \frac{\partial u_z}{\partial y} \right) \\ \mu \left( \frac{\partial u_z}{\partial x} + \frac{\partial u_x}{\partial z} \right) & \mu \left( \frac{\partial u_z}{\partial y} + \frac{\partial u_y}{\partial z} \right) & 2\mu \frac{\partial u_z}{\partial z} + \lambda \nabla \cdot \mathbf{v} \end{pmatrix} \quad (17)$$

where  $\mathbf{X} = [x \ y \ z]$  while  $\lambda$  is the second viscosity of the fluid, which correspond to the viscous effect due to compression or dilatation [45]. Mathematically,  $\lambda = 2\mu/3 - n'$ , where  $n'$  can be cancelled out when

the fluid exists as monatomic gas at low density [46]. When the flow is incompressible,  $\nabla \cdot \mathbf{v} = 0$  takes the second viscosity to be ignorable. By combining Eq. (16) and (17) will form:

$$\boldsymbol{\tau} = \langle \mu \nabla^2 u_x \rangle \mathbf{e}_x + \langle \mu \nabla^2 u_y \rangle \mathbf{e}_y + \langle \mu \nabla^2 u_z \rangle \mathbf{e}_z = \left\langle \mu \nabla^2 \mathbf{v} \middle| \nabla^2 = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right\rangle \quad (18)$$

With the presence of pressure gradient per unit volume, Eq. (15) and Eq. (18) can be incorporated to form the Navier-Stokes Equations as in Eq. (19).

$$\rho \frac{D\mathbf{v}}{Dt} = -\frac{\partial P}{\partial \mathbf{x}} + \mu \nabla^2 \mathbf{v} + \rho \mathbf{g}_x \quad (19)$$

### 2.2.2 Cylindrical coordinate Navier-Stokes Equations

The material derivative in cylindrical coordinate can be obtained as in Eq. (20).

$$\because u_r \partial t = \partial r, \quad u_\theta \partial t = r \partial \theta, \quad u_z \partial t = \partial z$$

$$\frac{D\mathbf{v}}{Dt} = \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{v}}{\partial r} \frac{\partial r}{\partial t} + \frac{\partial \mathbf{v}}{\partial \theta} \frac{\partial \theta}{\partial t} + \frac{\partial \mathbf{v}}{\partial z} \frac{\partial z}{\partial t} \Leftrightarrow \frac{\partial \mathbf{v}}{\partial t} + u_r \frac{\partial \mathbf{v}}{\partial r} + \frac{u_\theta}{r} \frac{\partial \mathbf{v}}{\partial \theta} + u_z \frac{\partial \mathbf{v}}{\partial z} \quad (20)$$

Expanding Eq. (20) with respect to its spatial components,

$$\frac{D\mathbf{v}}{Dt} = \frac{\partial(\mathbf{v} \cdot \mathbf{e}_x)}{\partial t} + \frac{u_\theta}{r} \frac{\partial(\mathbf{v} \cdot \mathbf{e}_x)}{\partial \theta} + u_z \frac{\partial(\mathbf{v} \cdot \mathbf{e}_x)}{\partial z}$$

Note that from differential operations in curvilinear coordinates [47,48],

$$\frac{\partial \mathbf{e}_r}{\partial \theta} = \mathbf{e}_\theta \quad \text{and} \quad \frac{\partial \mathbf{e}_\theta}{\partial \theta} = -\mathbf{e}_r$$

$$\therefore \frac{D\mathbf{v}}{Dt} = \left\langle \frac{\partial u_r}{\partial t} + u_r \frac{\partial u_r}{\partial r} + u_\theta \left( \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} \right) + \frac{\partial u_r}{\partial z} \right\rangle \mathbf{e}_r + \left\langle \frac{\partial u_\theta}{\partial t} + u_r \frac{\partial u_\theta}{\partial r} + u_\theta \left( \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} \right) + \frac{\partial u_\theta}{\partial z} \right\rangle \mathbf{e}_\theta + \left\langle \frac{\partial u_z}{\partial t} + u_r \frac{\partial u_z}{\partial r} + u_\theta \left( \frac{1}{r} \frac{\partial u_z}{\partial \theta} \right) + \frac{\partial u_z}{\partial z} \right\rangle \mathbf{e}_z \quad (21)$$

While the stress tensors for the cylindrical fluid elements are:

$$\boldsymbol{\kappa} = \begin{pmatrix} \kappa_{r,r} & \kappa_{\theta,r} & \kappa_{z,r} \\ \kappa_{r,\theta} & \kappa_{\theta,\theta} & \kappa_{z,\theta} \\ \kappa_{r,z} & \kappa_{\theta,z} & \kappa_{z,z} \end{pmatrix} = \begin{pmatrix} \frac{\tau_{rr}}{r} + \frac{\partial \tau_{rr}}{\partial r} & \frac{1}{r} \frac{\partial \tau_{r\theta}}{\partial \theta} & \frac{\partial \tau_{rz}}{\partial z} \\ \frac{\tau_{r\theta}}{r} + \frac{\partial \tau_{r\theta}}{\partial r} & \frac{1}{r} \frac{\partial \tau_{\theta\theta}}{\partial \theta} & \frac{\partial \tau_{z\theta}}{\partial z} \\ \frac{\tau_{rz}}{r} + \frac{\partial \tau_{rz}}{\partial r} & \frac{1}{r} \frac{\partial \tau_{\theta z}}{\partial \theta} & \frac{\partial \tau_{zz}}{\partial z} \end{pmatrix} = \frac{\partial \boldsymbol{\tau}}{\partial \mathbf{x}} \quad (22)$$

Note that  $\delta r^2 \approx 0$  during the simplification process. Due to its mathematical nature in curvilinear coordinates [44,45], the stress tensor will be:

$$\boldsymbol{\tau} = \left\langle \frac{\tau_{rr}}{r} + \frac{\partial \tau_{rr}}{\partial r} + \frac{1}{r} \frac{\partial \tau_{r\theta}}{\partial \theta} + \frac{\partial \tau_{rz}}{\partial z} - \frac{\tau_{\theta\theta}}{r} \right\rangle \mathbf{e}_r + \left\langle \frac{\tau_{r\theta}}{r} + \frac{\partial \tau_{r\theta}}{\partial r} + \frac{1}{r} \frac{\partial \tau_{\theta\theta}}{\partial \theta} + \frac{\partial \tau_{z\theta}}{\partial z} + \frac{\tau_{\theta r}}{r} \right\rangle \mathbf{e}_\theta + \left\langle \frac{\tau_{rz}}{r} + \frac{\partial \tau_{rz}}{\partial r} + \frac{1}{r} \frac{\partial \tau_{\theta z}}{\partial \theta} + \frac{\partial \tau_{zz}}{\partial z} \right\rangle \mathbf{e}_z \quad (23)$$

The single deformation rate,  $\zeta$  could be formulated from the material derivatives from Eq. (21) by removing the unsteady term and the velocity vectors prior to the operators, as shown in Eq. (24). The stress tensor,  $\boldsymbol{\tau}$  can be consequently transformed into the deformation rate tensor by the summation of the single deformation and inverse of single deformation.

$$\boldsymbol{\zeta} = \begin{pmatrix} \zeta_{rr} & \zeta_{r\theta} & \zeta_{rz} \\ \zeta_{\theta r} & \zeta_{\theta\theta} & \zeta_{\theta z} \\ \zeta_{zr} & \zeta_{z\theta} & \zeta_{zz} \end{pmatrix} = \begin{pmatrix} \frac{\partial u_r}{\partial r} & \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} & \frac{\partial u_r}{\partial z} \\ \frac{\partial u_\theta}{\partial r} & \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} & \frac{\partial u_\theta}{\partial z} \\ \frac{\partial u_z}{\partial r} & \frac{1}{r} \frac{\partial u_z}{\partial \theta} & \frac{\partial u_z}{\partial z} \end{pmatrix} \quad (24)$$

$$\boldsymbol{\tau} = \begin{pmatrix} \tau_{rr} & \tau_{r\theta} & \tau_{rz} \\ \tau_{\theta r} & \tau_{\theta\theta} & \tau_{\theta z} \\ \tau_{zr} & \tau_{z\theta} & \tau_{zz} \end{pmatrix} = \mu(\boldsymbol{\zeta} + \boldsymbol{\zeta}^{-1}) = \mu \begin{pmatrix} 2 \frac{\partial u_r}{\partial r} & \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} + \frac{\partial u_\theta}{\partial r} & \frac{\partial u_r}{\partial z} + \frac{\partial u_z}{\partial r} \\ \frac{\partial u_\theta}{\partial r} + \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} & 2 \left( \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} \right) & \frac{\partial u_\theta}{\partial z} + \frac{1}{r} \frac{\partial u_z}{\partial \theta} \\ \frac{\partial u_z}{\partial r} + \frac{\partial u_r}{\partial z} & \frac{1}{r} \frac{\partial u_z}{\partial \theta} + \frac{\partial u_\theta}{\partial z} & 2 \frac{\partial u_z}{\partial z} \end{pmatrix} \cdot (\mathbf{e}_r \quad \mathbf{e}_\theta \quad \mathbf{e}_z) \quad (25)$$

Incorporating Eq. (25) with Eq. (23) will form the cylindrical Navier-Stokes Equation.

$$\boldsymbol{\tau} = \left\langle \mu \left( \nabla^2 u_r - \frac{u_r}{r^2} - \frac{2}{r^2} \frac{\partial u_\theta}{\partial \theta} \right) \right\rangle \mathbf{e}_r + \left\langle \mu \left( \nabla^2 u_\theta - \frac{u_\theta}{r^2} + \frac{2}{r^2} \frac{\partial u_r}{\partial \theta} \right) \right\rangle \mathbf{e}_\theta + \left\langle \mu \nabla^2 u_z \right\rangle \mathbf{e}_z \left| \nabla^2 = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2}{\partial \theta^2} + \frac{\partial^2}{\partial z^2} \right. \quad (26)$$

$$\therefore \rho \frac{D\mathbf{v}}{Dt} = \frac{\partial \mathbf{P}}{\partial \mathbf{X}} + \boldsymbol{\tau} \boldsymbol{\chi} + \rho \mathbf{g}_X \quad (27)$$

### 2.2.3 Spherical coordinate Navier-Stokes Equations

The derivation procedure for spherical coordinate Navier-Stokes equations basically complies with all the steps as shown from Eq. (20) - (27). Therefore, only key equations and steps are unfolded here.

$$\therefore u_r \partial t = \partial r, u_\theta \partial t = r \partial \theta, u_\phi \partial t = r(\sin \theta) \partial \phi$$

$$\frac{D\mathbf{v}}{Dt} = \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{v}}{\partial r} \frac{\partial r}{\partial t} + \frac{\partial \mathbf{v}}{\partial \theta} \frac{\partial \theta}{\partial t} + \frac{\partial \mathbf{v}}{\partial \phi} \frac{\partial \phi}{\partial t} \Leftrightarrow \frac{\partial \mathbf{v}}{\partial t} + \mathbf{u}_r \frac{\partial \mathbf{v}}{\partial r} + \frac{u_\theta}{r} \frac{\partial \mathbf{v}}{\partial \theta} + \frac{u_\phi}{r \sin \theta} \frac{\partial \mathbf{v}}{\partial \phi} \quad (28)$$

The differential operations in curvilinear coordinates [47,48] for spherical domain are:

$$\frac{\partial \mathbf{e}_X}{\partial \mathbf{X}} = \begin{pmatrix} \frac{\partial \mathbf{e}_r}{\partial r} & \frac{\partial \mathbf{e}_\theta}{\partial r} & \frac{\partial \mathbf{e}_\phi}{\partial r} \\ \frac{\partial \mathbf{e}_r}{\partial \theta} & \frac{\partial \mathbf{e}_\theta}{\partial \theta} & \frac{\partial \mathbf{e}_\phi}{\partial \theta} \\ \frac{\partial \mathbf{e}_r}{\partial \phi} & \frac{\partial \mathbf{e}_\theta}{\partial \phi} & \frac{\partial \mathbf{e}_\phi}{\partial \phi} \end{pmatrix} = \begin{pmatrix} 0 & 0 & 0 \\ \mathbf{e}_\theta & -\mathbf{e}_r & 0 \\ \mathbf{e}_\phi \sin \theta & \mathbf{e}_\phi \cos \theta & -\mathbf{e}_r \sin \theta - \mathbf{e}_\theta \cos \theta \end{pmatrix} \quad (29)$$

$$\therefore \frac{D\mathbf{v}}{Dt} = \left\langle \frac{\partial u_r}{\partial t} + \mathbf{u}_r \frac{\partial u_r}{\partial r} + \mathbf{u}_\theta \left( \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} \right) + \mathbf{u}_\phi \left( \frac{1}{r \sin \theta} \frac{\partial u_r}{\partial \phi} - \frac{u_\phi}{r} \right) \right\rangle \mathbf{e}_r + \left\langle \frac{\partial u_\theta}{\partial t} + \mathbf{u}_r \frac{\partial u_\theta}{\partial r} + \mathbf{u}_\theta \left( \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} \right) + \mathbf{u}_\phi \left( \frac{1}{r \sin \theta} \frac{\partial u_\theta}{\partial \phi} - \frac{u_\phi \cot \theta}{r} \right) \right\rangle \mathbf{e}_\theta + \left\langle \frac{\partial u_\phi}{\partial t} + \mathbf{u}_r \frac{\partial u_\phi}{\partial r} + \frac{u_\theta}{r} \frac{\partial u_\phi}{\partial \theta} + \mathbf{u}_\phi \left( \frac{1}{r \sin \theta} \frac{\partial u_\phi}{\partial \phi} + \frac{u_r + u_\theta \cot \theta}{r} \right) \right\rangle \mathbf{e}_\phi \quad (30)$$

The stress tensor is formed based on the momentum conservation principle by referring to **Fig. 3**.

$$\boldsymbol{\tau} = \left\langle \frac{1}{r^2} \frac{\partial (\tau_{rr})}{\partial r} + \frac{1}{r \sin \theta} \frac{\partial (\tau_{\theta r} \sin \theta)}{\partial \theta} + \frac{1}{r \sin \theta} \frac{\partial \tau_{\phi r}}{\partial \phi} - \frac{\tau_{\theta\theta} + \tau_{\phi\phi}}{r} \right\rangle \mathbf{e}_r + \left\langle \frac{1}{r^3} \frac{\partial (r^3 \tau_{r\theta})}{\partial r} + \frac{1}{r \sin \theta} \frac{\partial (\tau_{\theta\theta} \sin \theta)}{\partial \theta} + \frac{1}{r \sin \theta} \frac{\partial \tau_{\phi\theta}}{\partial \phi} - \frac{\tau_{r\theta} + \tau_{\phi\theta} \cot \theta - \tau_{\theta r}}{r} \right\rangle \mathbf{e}_\theta + \left\langle \frac{1}{r^3} \frac{\partial (r^3 \tau_{r\phi})}{\partial r} + \frac{1}{r \sin \theta} \frac{\partial (\tau_{\theta\phi} \sin \theta)}{\partial \theta} + \frac{1}{r \sin \theta} \frac{\partial \tau_{\phi\phi}}{\partial \phi} - \frac{\tau_{r\phi} + \tau_{\phi\theta} \cot \theta + \tau_{\phi r}}{r} \right\rangle \mathbf{e}_z \quad (31)$$

The corresponding constitutive relationships will be:

$$\boldsymbol{\zeta} = \begin{pmatrix} \zeta_{rr} & \zeta_{r\theta} & \zeta_{rz} \\ \zeta_{\theta r} & \zeta_{\theta\theta} & \zeta_{\theta\phi} \\ \zeta_{\phi r} & \zeta_{\phi\theta} & \zeta_{\phi\phi} \end{pmatrix} = \begin{pmatrix} \frac{\partial u_r}{\partial r} & \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} & \frac{1}{r \sin \theta} \frac{\partial u_r}{\partial \phi} - \frac{u_\phi}{r} \\ \frac{\partial u_\theta}{\partial r} & \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} & \frac{1}{r \sin \theta} \frac{\partial u_\theta}{\partial \phi} - \frac{u_\phi \cot \theta}{r} \\ \frac{\partial u_\phi}{\partial r} & \frac{1}{r} \frac{\partial u_\phi}{\partial \theta} & \frac{1}{r \sin \theta} \frac{\partial u_\phi}{\partial \phi} + \frac{u_r + u_\theta \cot \theta}{r} \end{pmatrix} \quad (32)$$

$$\boldsymbol{\tau} = \begin{pmatrix} \tau_{rr} & \tau_{r\theta} & \tau_{r\phi} \\ \tau_{\theta r} & \tau_{\theta\theta} & \tau_{\theta\phi} \\ \tau_{\phi r} & \tau_{\phi\theta} & \tau_{\phi\phi} \end{pmatrix} = \mu \begin{pmatrix} 2 \frac{\partial u_r}{\partial r} & \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} + \frac{\partial u_\theta}{\partial r} & \frac{1}{r \sin \theta} \frac{\partial u_r}{\partial \phi} - \frac{u_\phi}{r} + \frac{\partial u_\phi}{\partial r} \\ \frac{\partial u_\theta}{\partial r} + \frac{1}{r} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{r} & 2 \left( \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} \right) & \frac{1}{r \sin \theta} \frac{\partial u_\theta}{\partial \phi} - \frac{u_\phi \cot \theta}{r} + \frac{1}{r} \frac{\partial u_\phi}{\partial \theta} \\ \frac{\partial u_\phi}{\partial r} + \frac{1}{r \sin \theta} \frac{\partial u_r}{\partial \phi} - \frac{u_\phi}{r} & \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{1}{r \sin \theta} \frac{\partial u_\theta}{\partial \phi} - \frac{u_\phi \cot \theta}{r} & 2 \left( \frac{1}{r \sin \theta} \frac{\partial u_\phi}{\partial \phi} + \frac{u_r + u_\theta \cot \theta}{r} \right) \end{pmatrix} \quad (33)$$

$$\boldsymbol{\tau} = \left\{ \begin{aligned} &\mu \left( \nabla^2 u_r - \frac{2u_r}{r^2} - \frac{2}{r^2} \frac{\partial u_\theta}{\partial \theta} - \frac{2u_\theta \cot \theta}{r^2} - \frac{2}{r^2 \sin \theta} \frac{\partial u_\phi}{\partial \phi} \right) \mathbf{e}_r + \mu \left( \nabla^2 u_\theta + \frac{2}{r^2} \frac{\partial u_r}{\partial \theta} - \frac{u_\theta}{(r \sin \theta)^2} - \frac{2 \cot \theta}{r^2 \sin \theta} \frac{\partial u_\phi}{\partial \phi} \right) \mathbf{e}_\theta \\ &+ \mu \left( \nabla^2 u_z + \frac{2}{r^2 \sin \theta} \frac{\partial u_r}{\partial \phi} - \frac{u_\phi}{(r \sin \theta)^2} - \frac{2 \cot \theta}{r^2 \sin \theta} \frac{\partial u_\theta}{\partial \phi} \right) \mathbf{e}_z \end{aligned} \right\} \left| \nabla^2 = \frac{1}{r^2} \frac{\partial^2 (r^2)}{\partial \theta^2} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right\} \quad (34)$$

When the pressure effect and body force effect step in, the spherical Navier-Stokes Equations will be formed in the similar way as in Eq. (27).

### 2.3 Energy Equations

The energy increment rate per unit volume,  $\Delta E$  consists of kinetic term and internal term, while the net heat flux going into the control volume per unit volume,  $\dot{q}$ . Referring to **Fig. 1**,

$$\Delta E = \rho \frac{D}{Dt} \left( e + \frac{1}{2} \mathbf{v}^2 \right) \quad (35)$$

$$\dot{q} = - \left( \frac{\partial q_x}{\partial x} + \frac{\partial q_y}{\partial y} + \frac{\partial q_z}{\partial z} \right) = \left\{ -\nabla \cdot \mathbf{q} \mid \nabla = \frac{\partial}{\partial x} + \frac{\partial}{\partial y} + \frac{\partial}{\partial z} \right\} \quad (36)$$

The work done on the control volume,  $\dot{W}$  is the product of wall shear stress or pressure, and the velocity component at a fluid element surface:

$$\dot{W} = \left\{ -\nabla \cdot (\mathbf{P}\mathbf{v}) + \frac{\partial (\boldsymbol{\tau}_X \mathbf{u}_X)}{\partial X} + \rho \mathbf{V} \mathbf{g} \mid \boldsymbol{\tau}_X \parallel \mathbf{X} \right\} \quad (37)$$

where  $\rho \mathbf{V} \mathbf{g}$  represents the work done by the volume force per unit mass which acts on the fluid such as gravity. The pressure and shear stress components are the effect of surface force while the work done by the moving fluid is analogue to the effect of body force as explained in the previous section.

Combining Eq. (35) - (37):

$$\rho \frac{D}{Dt} \left( e + \frac{1}{2} \mathbf{v}^2 \right) = -\nabla \cdot \mathbf{q} - \nabla \cdot (\mathbf{P}\mathbf{v}) + \frac{\partial (\boldsymbol{\tau}_X \mathbf{u}_X)}{\partial X} + \dot{Q}_{\text{volume}} + \rho \mathbf{V} \cdot \mathbf{g} \Leftrightarrow -\nabla \cdot \mathbf{q} - \nabla \cdot (\mathbf{P}\mathbf{v}) + \mathbf{v} \cdot \mathbf{f} + \mu \Phi + \dot{Q}_{\text{volume}} + \rho \mathbf{V} \mathbf{g} \quad (38)$$

$$\mathbf{f} = \left( \frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{yx}}{\partial y} + \frac{\partial \tau_{zx}}{\partial z} \right) \mathbf{e}_x + \left( \frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{zy}}{\partial z} \right) \mathbf{e}_y + \left( \frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{yz}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z} \right) \mathbf{e}_z \quad (39)$$

$$\mu \Phi = \left( \tau_{xx} \frac{\partial u_x}{\partial x} + \tau_{yy} \frac{\partial u_y}{\partial y} + \tau_{zz} \frac{\partial u_z}{\partial z} \right) + \tau_{xy} \left( \frac{\partial u_y}{\partial x} + \frac{\partial u_x}{\partial y} \right) + \tau_{yz} \left( \frac{\partial u_z}{\partial y} + \frac{\partial u_y}{\partial z} \right) + \tau_{zx} \left( \frac{\partial u_x}{\partial z} + \frac{\partial u_z}{\partial x} \right) \quad (40)$$

where  $\dot{Q}_{\text{volume}}$  represents a volumetric heat generation per unit volume. Eq. (38) is the one form of the energy equation. However, we can rewrite it to other forms using the Cauchy momentum equation as in Eq. (13). The Cauchy momentum equation can be rewritten as:

$$\begin{aligned} \rho \frac{D\mathbf{v}}{Dt} &= \rho \mathbf{g} - \nabla P + \mathbf{f} \\ \rho \mathbf{V} \frac{D\mathbf{v}}{Dt} &= \frac{\rho \mathbf{V} D\mathbf{v}^2}{Dt} = \rho \mathbf{V} \mathbf{g} - \mathbf{V} \cdot \nabla P + \mathbf{V} \cdot \mathbf{f} \end{aligned} \quad (41)$$

∴ Fourier's law,  $\mathbf{q} = -k \nabla T$

Subtracting Eq. (41) from Eq. (38),

$$\rho \frac{De}{Dt} = \nabla \cdot (k \nabla T) - P (\nabla \cdot \mathbf{v}) + \mu \Phi + \dot{Q}_{\text{volume}} \quad (42)$$

Eq. (42) is another form of the energy equation. Note that the  $\rho \mathbf{V} \mathbf{g}$  term is vanished in Eq. (42). However, the potential energy based on  $\rho \mathbf{V} \mathbf{g}$  is implicitly included in Eq. (42) since the potential energy is considered in Eq. (41). Continuity equation as in Eq. (2) can be rewritten as

$$\nabla \cdot \mathbf{v} = - \frac{1}{\rho} \frac{D\rho}{Dt} \quad (43)$$

By substituting Eq. (43) into (42), Eq. (41) becomes



$$\rho \left( \frac{De}{Dt} + P \frac{D\rho^{-1}}{Dt} \right) = \nabla \cdot (k \nabla T) + \mu \Phi + \dot{Q}_{\text{volume}}$$

$$h = e + P/\rho$$

$$\rho \frac{Dh}{Dt} = \frac{DP}{Dt} + \nabla \cdot (k \nabla T) + \mu \Phi + \dot{Q}_{\text{volume}} \quad (44)$$

Eq. (44) is the most common form of the energy equation. Expanding viscous tensor,  $\Phi$ :

$$\Phi = \begin{cases} \left[ \left( \frac{\partial u_x}{\partial x} + \frac{\partial u_x}{\partial y} \right)^2 + \left( \frac{\partial u_z}{\partial y} + \frac{\partial u_y}{\partial z} \right)^2 + \left( \frac{\partial u_x}{\partial z} + \frac{\partial u_z}{\partial x} \right)^2 + \frac{2}{3} \left( \left( \frac{\partial u_x}{\partial x} - \frac{\partial u_y}{\partial y} \right)^2 + \left( \frac{\partial u_y}{\partial y} - \frac{\partial u_z}{\partial z} \right)^2 + \left( \frac{\partial u_z}{\partial z} - \frac{\partial u_x}{\partial x} \right)^2 \right) \right] \text{ Cartesian} \\ \left( \frac{\partial u_\theta}{\partial x} - \frac{u_\theta}{r} + \frac{1}{r} \frac{\partial u_r}{\partial \theta} \right)^2 + \left( \frac{1}{r} \frac{\partial u_z}{\partial \theta} + \frac{\partial u_\theta}{\partial z} \right)^2 + \left( \frac{\partial u_r}{\partial z} + \frac{\partial u_z}{\partial r} \right)^2 + 2 \left( \left( \frac{\partial u_r}{\partial r} \right)^2 + \left( \frac{1}{r} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} \right)^2 + \left( \frac{\partial u_z}{\partial z} \right)^2 \right) - \frac{2}{3} (\nabla \cdot \mathbf{v})^2 \text{ Cylindrical} \\ \left( r \frac{\partial}{\partial r} \left( \frac{u_\phi}{r} \right) + \frac{1}{r} \frac{\partial u_r}{\partial \phi} \right)^2 + \left( \frac{\sin \phi}{r} \frac{\partial}{\partial \theta} \left( \frac{u_\theta}{r \sin \phi} \right) + \frac{1}{r \sin \phi} \frac{\partial u_\theta}{\partial \theta} \right)^2 + \left( \frac{1}{r \sin \phi} \frac{\partial u_r}{\partial \theta} + r \frac{\partial}{\partial r} \left( \frac{u_\theta}{r} \right) \right)^2 + 2 \left( \left( \frac{\partial u_r}{\partial r} \right)^2 + \left( \frac{1}{r} \frac{\partial u_\phi}{\partial \phi} + \frac{u_r}{r} \right)^2 + \left( \frac{1}{r \sin \phi} \frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r} + \frac{u_\theta \cot \phi}{r} \right)^2 \right) - \frac{2}{3} (\nabla \cdot \mathbf{v})^2 \text{ Spherical} \end{cases} \quad (45)$$

The specific enthalpy is a function of temperature and pressure. It can be expressed as

$$dh = \left( \frac{\partial h}{\partial T} \right)_P dT + \left( \frac{\partial h}{\partial P} \right)_T dP \quad (46)$$

The thermodynamic relations give us

$$\left( \frac{\partial h}{\partial T} \right)_P = C_P, \quad \left( \frac{\partial h}{\partial P} \right)_T = \frac{1 - \beta T}{\rho} \quad (47)$$

where  $C_P$  is the specific heat at constant pressure and  $\beta$  is the volume expansivity,  $\beta = -\frac{1}{\rho} \left( \frac{\partial \rho}{\partial V} \right)_P$ . If the fluid is an ideal gas, the volume expansivity is

$$\beta = \frac{1}{T} \quad (48)$$

By substituting Eq. (48) into Eq. (47), the specific enthalpy is expressed as

$$dh = C_P dT \quad (49)$$

By substituting Eq. (49) into Eq. (44), the energy equation for an ideal gas is obtained as:

$$\rho C_P \frac{DT}{Dt} = \frac{DP}{Dt} + \nabla \cdot (k \nabla T) + \mu \Phi + \dot{Q}_{\text{volume}} \quad (50)$$

If the density of a fluid is constant. The volume expansion of such a constant density fluid is zero:

$$dh = C_P dT + \frac{1}{\rho} dP \quad (51)$$

By substituting Eq. (51) into Eq. (44),

$$\rho C_P \frac{DT}{Dt} = \nabla \cdot (k \nabla T) + \mu \Phi + \dot{Q}_{\text{volume}} \quad (52)$$

Let's consider a steady ideal gas flow without volume heat generation whose properties are constant except the density. The Cartesian coordinate system is considered for convenience. Using the following dimensionless variables:

$$X = \frac{x}{L}, \quad Y = \frac{y}{L}, \quad Z = \frac{z}{L}, \quad U_x = \frac{u_x}{u_{ref}}, \quad U_y = \frac{u_y}{u_{ref}}, \quad U_z = \frac{u_z}{u_{ref}},$$

$$P = \frac{P}{\rho_{ref} u_{ref}^2}, \quad \Theta = \frac{T}{T_{ref}}, \quad Re = \frac{\rho_{ref} u_{ref} L}{\mu}, \quad Ma = \frac{u_{ref}}{\sqrt{\gamma R_{gas} T_{ref}}}, \quad \rho^* = \frac{\rho}{\rho_{ref}}$$

The energy equation is expressed as

$$\rho^* \left( U_X \frac{\partial \Theta}{\partial X} + U_Y \frac{\partial \Theta}{\partial Y} + U_Z \frac{\partial \Theta}{\partial Z} \right) = (\gamma - 1) \text{Ma}^2 \left( U_X \frac{\partial P}{\partial X} + U_Y \frac{\partial P}{\partial Y} + U_Z \frac{\partial P}{\partial Z} \right) + \frac{1}{\text{RePr}} \left[ \frac{\partial^2 \Theta}{\partial X^2} + \frac{\partial^2 \Theta}{\partial Y^2} + \frac{\partial^2 \Theta}{\partial Z^2} \right] + (\gamma - 1) \frac{\text{Ma}^2}{\text{Re}} \Phi^* \quad (53)$$

where  $\Phi^*$  is the dimensionless viscous dissipation function and it can be expressed as:

$$\Phi^* = \left( \left( \frac{\partial U_Y}{\partial X} + \frac{\partial U_X}{\partial Y} \right)^2 + \left( \frac{\partial U_Z}{\partial Y} + \frac{\partial U_Y}{\partial Z} \right)^2 + \left( \frac{\partial U_X}{\partial Z} + \frac{\partial U_Z}{\partial X} \right)^2 \right) + \frac{2}{3} \left( \left( \frac{\partial U_X}{\partial X} - \frac{\partial U_Y}{\partial Y} \right)^2 + \left( \frac{\partial U_Y}{\partial Y} - \frac{\partial U_Z}{\partial Z} \right)^2 + \left( \frac{\partial U_Z}{\partial Z} - \frac{\partial U_X}{\partial X} \right)^2 \right) \quad (54)$$

Note that the first and the third terms of right hand side of Eq. (54) are the substantial derivative of pressure term (SDP) and the viscous dissipation term (VD), respectively.  $\text{Ma}^2$  is multiplied to both terms. Therefore, the both SDP and VD terms can be neglected when the Ma of the flow is less than 0.3. However, the both terms should be neglected simultaneously. If one of the both terms remains, this results in physically unrealistic result [49]. Then, the energy equation of an ideal gas flow with low velocity becomes:

$$\rho C_p \frac{DT}{Dt} = \nabla \cdot (k \nabla T) + \dot{Q}_{\text{volume}} \quad (55)$$

The same form of the equation is obtained for a constant density fluid flow if the viscous dissipation is negligible. Usually, the viscous dissipation term for a liquid flow in a conventional size tube is negligible. However, the velocity gradient of a flow in a small sized tube becomes huge, it is to say that the viscous dissipation of a liquid flow in a micro tube whose diameter is less than 200  $\mu\text{m}$  is not negligible [50, 51]. In such a case, Eq. (52) should be solved.

### 3. Developments and Future Research of CFD Governing Equations

Navier-Stokes equations are the pillar for all the fluid flow dynamics [35], with very wide applications in engineering such as aerodynamics [52-54], fluid-structure interaction [55-60], turbomachinery [61-64], biomedical simulations [65-68], nanofluids [10,69-71] and bio-inspired transportation [72-75]. Due to its omnipresent applications, providing the solution to the Navier-Stokes Equations becomes one of the largest interests among the researchers. Due to its mathematical perplexity and its application-wise complexity, the equations are solved numerically in three ways: fixed-grid methods [36,40,76-81], immersed boundary methods [82-84], meshfree methods [85-90] and other numerical schemes such as Runge-Kutta method [91,92].

However, Navier-Stokes Equations do not consider many other factors and researches are working on to complement them. Dong and Wu [93] claimed that the current Navier-Stokes Equations have under-estimated the fluid forces, as the rotations, changes in shear rate and turbulence [94] will add in more forces. They modified Eq. (19) to be:

$$\rho \frac{D\mathbf{v}}{Dt} = -\frac{\partial P}{\partial \mathbf{X}} + \mu \nabla^2 \mathbf{v} + \rho \mathbf{g}_X + f_u^A + f_u^B + f_u^R + f_u^S \quad (56)$$

where  $f_u^A$ ,  $f_u^B$ ,  $f_u^R$  and  $f_u^S$  is additional unsteady force with  $u$ , additional history force with  $u$ , additional rotational force with  $u$  and additional gradient force, respectively.

The flow is also sometimes associated with a source of vorticity such as pressure gradient due to non-slip boundary conditions and Coriolis effect. Such a flow is named as non-stationary. Ershkov [95,96] modified the Navier-Stokes Equations and developed an approximate solution for it using Riccati partial differential equation [97].

$$0 = -\frac{\partial P}{\partial \mathbf{X}} + \mu \nabla^2 \mathbf{v} + \frac{1}{2} (\overrightarrow{u_p} + \overrightarrow{u_w})^2 \quad (57)$$

$\overrightarrow{u_p}$  is the irrotational field of flow while  $\overrightarrow{u_w}$  is a solenoidal field of flow velocity which generates a curl field. When the flow becomes turbulent, flow field fluctuation set in and more unknowns

transpires. This calls for the necessity to model more equations in order to close the existing equations. Navier-Stokes equations will need to be modified as Reynolds-Averaged Navier-Stokes (RANS) Equations [98,99] as shown in Eq. (59), and therefore turbulence modelling [100-104] is then introduced. Large eddy simulation [105-107] and direct numerical simulation [108-110] are alternatives too to deal with RANS equations.

$$\rho \frac{D\bar{\mathbf{v}}}{Dt} = -\frac{\partial \bar{P}}{\partial \mathbf{X}} + \mu \nabla^2 \bar{\mathbf{v}} + \rho \mathbf{g}_X - \rho \left( \frac{\partial (\bar{\mathbf{v}}', \bar{\mathbf{v}}')}{\partial \mathbf{X}} \right) \quad (58)$$

Note that  $\bar{P}$  and  $\bar{\mathbf{v}}$  is the average pressure and velocity respectively, while  $\bar{\mathbf{v}}'$  represents the velocity fluctuation. The term  $\rho \left( \frac{\partial (\bar{\mathbf{v}}', \bar{\mathbf{v}}')}{\partial \mathbf{X}} \right)$  is named as Reynolds stress. RANS equations will be significant too at a distance far away from the non-slip wall, as the eddies and turbulence production is high. When the speed of fluid goes beyond unity Mach number, the Navier-Stokes equations need to be modified to fit with the compressible flow [111-113], whereby the second viscosity in Eq. (17) can't be ignored as  $\nabla \mathbf{v} \neq 0$ .

Meanwhile energy equations are widely applied in computational heat transfer, which is often coupled with the computational fluid dynamics where the Continuity equations and Navier-Stokes equations are applied. Heat transfer analysis for internal flow [114-117], turbo-machinery [119-121] and biological heat transfer [122,123] are amongst the key research areas which needs energy equation as its governing equations.

In short, the development of both Navier-Stokes equations and energy equations move towards the inclusion of more other boundary factors, which re-conciliate the experimental and numerical techniques with the existing equations. Establishment of such ansatz may simplify the procedures in numerical analysis with the reduction of complexity in boundary condition treatment. The future development of the fluid dynamics equations, is not only confined to the solution of the equations using various techniques, but also the improvement of the equations which could predict real phenomenon more accurately. The recommendations for the future development could comprise the development of non-Newtonian momentum equations, formation of conservation equations in advanced coordinate systems and inclusion of more body forces into momentum equations.

### 3.1 Development of non-Newtonian Momentum Equations

Incorporation and simplification of the viscous term for non-Newtonian fluid [124,125]. Although there are a quite number of rheological correlation between the shear stress and viscosity [126-129], the incorporation of the non-Newtonian term into the Navier-Stokes and energy equations will expedite the modelling. This will expand the research horizon into the numerical and algorithm development for solving the engineering problems through computational rheology [130,131].

In this paper, the Ostwald-de Waele power law [132] is taken as the example to be incorporated into momentum equations. By considering the Cartesian coordinate, Eq. (17) can be modified as:

$$\boldsymbol{\tau} = \begin{pmatrix} \mu \left( 2 \frac{\partial u_x}{\partial x} \right)^n + \lambda \nabla \cdot \mathbf{v} & \mu \left( \frac{\partial u_x}{\partial y} + \frac{\partial u_y}{\partial x} \right)^n & \mu \left( \frac{\partial u_x}{\partial z} + \frac{\partial u_z}{\partial x} \right)^n \\ \mu \left( \frac{\partial u_y}{\partial x} + \frac{\partial u_x}{\partial y} \right)^n & \mu \left( 2 \frac{\partial u_y}{\partial y} \right)^n + \lambda \nabla \cdot \mathbf{v} & \mu \left( \frac{\partial u_y}{\partial z} + \frac{\partial u_z}{\partial y} \right)^n \\ \mu \left( \frac{\partial u_z}{\partial x} + \frac{\partial u_x}{\partial z} \right)^n & \mu \left( \frac{\partial u_z}{\partial y} + \frac{\partial u_y}{\partial z} \right)^n & \mu \left( 2 \frac{\partial u_z}{\partial z} \right)^n + \lambda \nabla \cdot \mathbf{v} \end{pmatrix} \quad (59)$$

Substituting Eq. (60) into (19) will lead to Eq. (61):

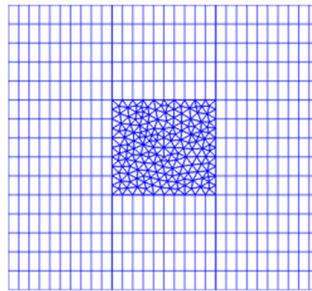
$$\rho \frac{Du_{\mathbf{X}}}{Dt} = -\frac{\partial P}{\partial \mathbf{X}} + \mu \frac{\partial}{\partial \mathbf{X}} \left( \frac{\partial u_{x_j}}{\partial x_i} + \frac{\partial u_{x_i}}{\partial x_j} \right)^n + \rho \mathbf{g}_X \quad (60)$$

where  $i$  and  $j$  is the spatial component, based on the stress tensor as in Eq. (59).  $n$  is the flow behaviour index. If  $n < 1$  and  $n > 1$  the fluids will exhibit pseudoplastic and dilatant behaviour respectively. The typical example for the former and later fluids is blood plasma and corn starch in ethylene glycol [133]. Eq. (61) can be therefore be applied to solve through various numerical methods such as SIMPLE algorithm [36] and stream-vorticity functions [76].

There are quite a number of non-Newtonian models, which can be found in several textbooks [41,124,125,134-136], which can be indeed further incorporated into the momentum equations. Furthermore, the expansion of the equations into curvilinear coordinate will open up a room for further researches. Such incorporation will path the way for the numerical simulation on non-Newtonian fluids and its conciliation with experimental results. Such development will bridge the gap that many non-Newtonian fluids applications such as blood flow [137-139], food processing [140-142] and chemical reactors [143-145] are simulated based the Stokes Law most of the time.

### 3.2 Formation of Conservation Equations in Advanced Coordinate Systems

Several meshing techniques are proposed in order to deal with the meshing issue. One of them is hybrid meshing [146] which combines the structured, unstructured and chimera grids in order to conform with the problem domain, as shown in **Fig. 4**. Such hybrid meshing calls for great energy when the boundary is in curvature form, while poses potential errors that lead to inconsistency and inaccuracy, including improper skewness, wrap angle ad aspect ratio. Local refinement at the boundary is possible, yet it will consume heavier computational effort.



**Fig. 4.** Hybrid meshing [147]

Some other more sophisticated methods of meshing are proposed too such as medial axis transform [148], all-quad meshing [149], dual contouring tetrahedral decomposition [150], high order curvilinear meshing [151] and radial basis function mesh deformation [152].

To enhance the room for computation without the cost of complex meshing, the Continuity equation, Navier-Stokes equations and energy equations in more advanced coordinate system can be developed. The possible system of coordinate could comprise ellipsoid, cone, hyperboloid and elliptical-paraboloid coordinate. The coordinate equation for ellipsoid, cone, hyperboloid and elliptical-paraboloid is from Eq. (61.1) - (61.4) respectively.

$$\left(\frac{x}{a}\right)^2 + \left(\frac{y}{b}\right)^2 + \left(\frac{z}{c}\right)^2 = 1 \quad (61.1)$$

$$\left(\frac{x}{a}\right)^2 + \left(\frac{y}{b}\right)^2 - \left(\frac{z}{c}\right)^2 = 0 \quad (61.2)$$

$$\left(\frac{x}{a}\right)^2 + \left(\frac{y}{b}\right)^2 - \left(\frac{z}{c}\right)^2 = 1 \quad (61.3)$$

$$\left(\frac{x}{a}\right)^2 + \left(\frac{y}{b}\right)^2 - \frac{z}{c} = 0 \quad (61.4)$$

$a$ ,  $b$  and  $c$  is the radius in the  $x$ -,  $y$ - and  $z$ - component respectively. However, the main problems in introducing the curvilinear conservation equations are: (1) mathematical complexity of determining the spatial derivatives during the derivations; and (2) singularity issue when the radius is approaching zero. To prevent the singularity issue, sometimes the curvilinear coordinate equations are not applied, while Cartesian equations will take place with some refinement on the meshing [148-152].

However, this remains a great field to be proceeded to enable to more robust computations without meshing issues for quite a wide range of engineering simulations such as airfoils investigation, shipping hydrodynamics, turbo-machineries flow dynamics and processing reactors simulations.

### 3.3 Inclusion of More Body Forces into Momentum Equations

Addition of more body force components into conservative equations, especially the momentum equations will be the next possible research direction. Most of the time, only gravitational force is considered during the modeling, as shown in Eq. (19) and (37). The introduction of electromagnetic force, Coriolis force, centrifugal force, impact force and vibrating force into the conservative equations will complement the equations, which will bring the CFD research to a greater height in investigating engineering physics. In adherence with the dimensional homogeneity of Eq. (19), the modified Navier-Stokes equations with the inclusion of expression of various body force can be:

$$\frac{Du_x}{Dt} = -\frac{\partial P}{\partial X} + \nu \nabla^2 u_x + \rho \left( g_x + \frac{\bar{v}^2}{2} \right) + \frac{q}{V} (\vec{E} + \vec{v} \vec{H}) + 2\rho v_t \omega + \rho r \omega^2 + \frac{F_0}{V} \cos(\omega t) \quad (62)$$

where  $\bar{v}$ ,  $V$ ,  $q$ ,  $\vec{E}$ ,  $\vec{v}$ ,  $\vec{H}$ ,  $v_t$ ,  $\omega$ ,  $F_0$  and  $\nu$  is vector moving speed of the control volume, volume of infinitesimal fluid element ( $m^3$ ), electric charge (Coulombs), electric field (volts per meter), charge velocity (m/s), magnetic field (Tesla), tangential velocity (m/s), radial speed (rad/s), vibrating force (Newton) and kinematic viscosity ( $m^2/s$ ) respectively.

Do note that the term  $\rho \frac{\bar{v}^2}{2}$  represents impact force,  $\frac{q}{V} (\vec{E} + \vec{v} \vec{H})$  is named as Lorentz force equation which represent the electromagnetic force [153],  $2\rho v_t \omega + \rho r \omega^2$  is the Coriolis force or centrifugal force [154] while  $\frac{F_0}{V} \cos(\omega t)$  denotes vibrating effects.

It is noteworthy that due to the complexity in the mechanical vibration,  $F_0$  is subjected to various parameters such as forced vibration, free vibration and damping. This will introduce fluid structure interaction with, in which such inclusion of body force may negate the necessity to implement immersed boundary methods [155] and meshfree methods [156,157] in the numerical solution.

## 4. Conclusion

The full derivations of Continuity equations, Navier-Stokes equations and energy equations have been structured out with physical explanation, applications and development. Recommendations on the future research in the development of governing equations are illustrated too. Development of non-Newtonian momentum equations, introduction of advanced system of coordinate into conservation equations and incorporation of body forces into the momentum equations will help to bypass the necessity to apply complicated numerical techniques in solving fluid dynamics problems. This will suggest a new path for CFD research in the nearest future.

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